

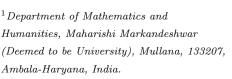
Curvature properties of α -cosymplectic manifolds with *- η -Ricci-Yamabe solitons

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ABSTRACT

In this research article, we study *- η -Ricci-Yamabe solitons on an α -cosymplectic manifold by giving an example in the support and also prove that it is an η -Einstein manifold. In addition, we investigate an α -cosymplectic manifold admitting *- η -Ricci-Yamabe solitons under some conditions. Lastly, we discuss the concircular, conformal, conharmonic, and W_2 -curvatures on the said manifold admitting *- η -Ricci-Yamabe solitons.

RESUMEN

En el presente artículo, estudiamos solitones *- η -Ricci-Yamabe en una variedad α -cosimpléctica dando un ejemplo que lo soporta y también probamos que es una variedad η -Einstein. Adicionalmente, investigamos una variedad α -cosimpléctica que admite solitones *- η -Ricci-Yamabe bajo ciertas condiciones. Finalmente, discutimos las curvaturas concircular, conforme, con-armónica y W_2 en dicha variedad admitiendo solitones *- η -Ricci-Yamabe.

Keywords and Phrases: *- η -Ricci-Yamabe soliton, α -cosymplectic manifold, curvature, η -Einstein manifold.

2020 AMS Mathematics Subject Classification: 53B20, 53C21, 53C44, 53C25, 53C50, 53D35

Published: 06 April, 2024 Accepted: 21 January, 2024 Received: 02 August, 2023





1 Introduction

In the year 1982, R. S. Hamilton [9] investigated the concept of Ricci flow on a smooth Riemannian manifold (shortly, RM). A self-similar solution to the Ricci flow is nothing but a Ricci soliton if it moves only by a one parameter family of diffeomorphism and scaling. After introducing the idea of Ricci flow, the theory of Yamabe flow was also initiated by Hamilton in [10] to construct Yamabe metrics on a compact RM. A Yamabe soliton is again corresponded to a self-similar solution of the Yamabe flow.

S. Guler and M. Crasmareanu gave a new class of geometric flow of type (ρ, q) , known as Ricci-Yamabe flow in [7]. They proposed the idea of Ricci-Yamabe soliton (shortly, RYS) if it moves only by one parameter group of diffeomorphism and scaling. The metric of the RM (M^n, h) , n > 2, is said to be RYS (h, V, Λ, ρ, q) if it satisfies the following [20]:

$$\pounds_V h + 2\rho Ric = [2\Lambda - qr]h, \tag{1.1}$$

where Lie derivative operator of the metric h along the vector field V represented by $\pounds_V h$, the Ricci curvature tensor by Ric (the Ricci operator Q defined by Ric(A, B) = h(QA, B) for $A, B \in \chi(M), \chi(M)$ being the Lie algebra of vector fields of M), the scalar curvature by r and the real scalars by Λ, ρ, q . According to Λ , RYS will be expanding, steady or shrinking if Λ is negative, zero or positive, respectively.

The concept of η -Ricci-Yamabe solitons (η -RYS) was defined by M. D. Siddiqi, et al. [20] in 2020 as a new generalization of RYS and it is defined as

$$\pounds_V h + 2Ric + [2\Lambda - qr]h + 2\mu\eta \otimes \eta = 0, \tag{1.2}$$

where μ is a constant and η is a 1-form on M.

On the other hand, S. Dey and S. Roy [5] inaugurated a new generalization of η -Ricci soliton (η -RS) [3], namely *- η -Ricci soliton (*- η -RS), defined below:

$$\pounds_V h + 2Ric^* + 2\Lambda h + 2\mu\eta \otimes \eta = 0, \tag{1.3}$$

where *-Ricci tensor (shortly, *-RT) is denoted by Ric^* .

Tachibana [22] brought up the concept of *-RT on almost Hermitian manifolds and afterwards Hamada [8] studied *-RT on real hypersurfaces of non-flat complex space forms. Such geometrical works inspired S. Roy, et al. to come up with new idea *- η -Ricci-Yamabe soliton (shortly, *- η -RYS) of type (ρ, q) , which is RM and fulfilling [18]

$$\pounds_V h + 2\rho Ric^* + [2\Lambda - qr^*]h + 2\mu\eta \otimes \eta = 0, \tag{1.4}$$

where $r^*(=trace(Ric^*))$ is the *-scalar curvature and Λ, ρ, q, μ are real scalars. The *- η -RYS is shrinking, steady or expanding if Λ is negative, zero or positive respectively. And they discussed *- η -RYS on α -cosymplectic manifolds with a quarter-symmetric metric (shortly, QSM) connection.

Further, A. Haseeb, R. Prasad and F. Mofarreh [12] obtained some interesting results on an α -Sasakian manifold admitting *- η -RYS with the potential vector field ζ satisfying conditions $Rim(\zeta, X).Ric = 0$, Q(h, Ric) = 0 and pseudo-Ricci symmetric and also showed that α -Sasakian admitting *- η -RYS is an η -EM.

In last few years, numerous authors have worked on the characterizations of Ricci, Ricci-Yamabe, η -Ricci-Yamabe and *- η -Ricci-Yamabe solitons (respectively, RS, RYS, η -RYS and *- η -RYS) in contact geometry. First, the study of RS in contact geometry was proposed by Sharma in [19]. After the initial work on Ricci solitons, some notable classes of contact geometry explored by H. I. Yoldas in [25,26] where Ricci solitons have been investigated. Later on, D. Dey [2] provided the idea of an almost Kenmotsu metric as RYS. Also, P. Zhang et al. [27] have studied conformal RYS on perfect fluid space-time. New type of soliton namely *-RYS on contact geometry introduced by M. D. Siddiqi and Akyol in [20] and they have discussed the notion of η -RYS for geometrical structure on Riemannian submersions admitting η -RYS with the potential field. In recent years, a Kenmotsu metric in terms of η -RYS was measured by Yoldas in [23]. Next, the notion of *- η -RYS was studied by many authors on different odd dimensional Riemannian manifolds. It should be noted that the geometry of *-k-RYS and gradient *-k-RYS on Kenmotsu manifolds were given by S. Dey and S. Roy in [4].

We organize this paper as follows: In section 2, we review some basic definitions and tools of an α -cosymplectic manifold M. The main results are stated in section 3. In fact, we prove that an n-dimensional M admitting a *- η -RYS is an η -Einstein manifold. Then some curvature tensor conditions are studied on M with *- η -RYS. Finally, in section 4, we discuss some results on M when it is ζ -concircularly flat, ζ -conharmonically flat, ζ -W₂ flat and ζ -conformal.

2 Preliminaries

On an n = 2m + 1-dimensional RM M, if an almost contact metric structure (Φ, ζ, η, h) satisfies the following relations, then M is called an almost contact metric manifold:

$$\Phi^2 A = A - \eta(A)\zeta \tag{2.1}$$

$$\eta(\zeta) = 1, \quad \Phi(\zeta) = 0, \quad \eta(\Phi\zeta) = 0$$
(2.2)



$$h(A, \Phi B) = -h(\Phi A, B), \tag{2.3}$$

$$h(A,\zeta) = \eta(A), \quad h(\Phi A, \Phi B) = h(A,B) - \eta(A)\eta(B), \tag{2.4}$$

for all $A, B \in \chi(M)$, where Φ denotes a (1,1) tensor field, ζ is a vector field, η is a 1-form and h is the compatible Riemannian metric.

The fundamental form ϕ on M is defined as [1]

$$\phi(A,B) = h(\Phi A, B), \tag{2.5}$$

for all $A, B \in \chi(M)$.

If the Nijenhuis tensor field of Φ on M satisfies $N_{\Phi}(A, B) + 2d\eta(A, B)\zeta = 0$, then M is called a normal almost contact metric manifold. Here

$$N_{\Phi}(A, B) = \Phi^{2}[A, B] + [\Phi A, \Phi B] - \Phi[A, \Phi B] - \Phi[\Phi A, B],$$

for any $A, B \in \chi(M)$.

Under the following conditions, a normal almost contact metric manifold M is known as an α -cosymplectic manifold (shortly, α -CM):

- (1) $d\eta = 0$,
- (2) $d\phi = 2\alpha\eta \wedge \phi$,

for $\alpha \in \mathbb{R}$.

We note that an α -CM can be

- (1) a cosymplectic manifold provided that $\alpha = 0$,
- (2) an α -Kenmotsu manifold provided that $\alpha \neq 0$.

For an α -CM M, we have

$$(\nabla_A \Phi) B = \alpha(h(\Phi A, B)\zeta - \eta(B)\Phi A) \tag{2.6}$$

and

$$\nabla_A \zeta = -\alpha \Phi^2 A = \alpha [A - \eta(A)\zeta], \tag{2.7}$$

where ∇ is the Levi-Civita connection associated with h.

The main examples and curvature characteristics of α -CM were firstly obtained in [11,14,15]. Also, we have the following relations for the Riemannian curvature tensor Rim and the Ricci curvature



tensor Ric of M:

$$Rim(A, B)\zeta = \alpha^2 \left[\eta(A)B - \eta(B)A \right], \tag{2.8}$$

$$Rim(\zeta, A)B = \alpha^2 \left[\eta(B)A - h(A, B)\zeta \right], \tag{2.9}$$

$$Rim(\zeta, A)\zeta = \alpha^2 \left[A - \eta(A)\zeta \right], \tag{2.10}$$

$$\eta(Rim(A, B)C) = \alpha^{2} [\eta(B)h(A, C) - \eta(A)h(B, C)], \qquad (2.11)$$

$$Ric(A,\zeta) = -\alpha^2(n-1)\eta(A), \tag{2.12}$$

for all $A, B, C \in \chi(M)$.

In [11], the *-RT Ric^* of type (0,2) on an n-dimensional α -CM M is obtained as

$$Ric^*(B, C) = Ric(B, C) + \alpha^2(n-2)h(B, C) + \alpha^2\eta(B)\eta(C),$$
 (2.13)

for any $B, C \in \chi(M)$.

Let $\{E_i|i=1,2,\ldots,n\}$ be an orthonormal basis of $T_p(M)$, $p \in M$. We set $B=C=E_i$ and it is easy to derive the *-scalar curvature $r^*=trace(Ric^*)$ as

$$r^* = r + \alpha^2 (n-1)^2. \tag{2.14}$$

On the other hand, α -CM M is said to be an η -EM if the Ricci curvature tensor has the following form [24]:

$$Ric(A, B) = uh(A, B) + v\eta(A)\eta(B), \tag{2.15}$$

for $A, B \in \chi(M)$, where u and v being constants.

For this paper, we need some curvature tensors on a RM (M^n, h) , which are given below [17]:

$$\overline{\mathbb{C}}(A,B)C = Rim(A,B)C - \frac{r}{n(n-1)}[h(B,C)A - h(A,C)B], \tag{2.16}$$

$$H(A,B)C = Rim(A,B)C - \frac{1}{n-2}[h(B,C)QA - h(A,C)QB + Ric(B,C)A - Ric(A,C)B],$$
(2.17)

$$W_2(A,B)C = Rim(A,B)C + \frac{1}{n-1}[h(A,C)QB - h(B,C)QA],$$
(2.18)

$$\mathbb{C}^{*}(A,B)C = Rim(A,B)C - \frac{1}{n-2}[Ric(B,C)A - Ric(A,C)B + h(B,C)QA - h(A,C)QB] + \frac{r}{(n-1)(n-2)}[h(B,C)A - h(A,C)B],$$
(2.19)

where $\overline{\mathbb{C}}$, H, W_2 and \mathbb{C}^* represent the concircular curvature tensor [16], the conharmonic curvature tensor [13], the W_2 -curvature tensor [16] and the conformal curvature tensor [6].



3 On α -CM admitting *- η -RYS

Let us take a *- η -RYS $(h, \zeta, \Lambda, \mu, \rho, q)$ on an n-dimensional α -CM M, which is given by

$$(\mathcal{L}_{\zeta}h)(A,B) + 2\rho Ric^*(A,B) + [2\Lambda - qr^*]h(A,B) + 2\mu\eta(A)\eta(B) = 0, \tag{3.1}$$

for any $A, B \in \chi(M)$.

Theorem 3.1. An n-dimensional α -CM M admitting *- η -RYS $(h, \zeta, \Lambda, \mu, \rho, q)$ is an η -EM of the constant scalar curvature r. Moreover, the scalars Λ and μ are related by

$$\Lambda + \mu = \frac{qr}{2} + \frac{q\alpha^2(n-1)^2}{2}.$$
 (3.2)

Proof. From (2.4) and (2.7), we arrive at

$$(\pounds_{\zeta}h)(A,B) = h(\nabla_A\zeta,B) + h(A,\nabla_B\zeta) = 2\alpha \bigg(h(A,B) - \eta(A)\eta(B)\bigg). \tag{3.3}$$

Substitute (3.3)) into (3.1) to get

$$Ric^*(A,B) = -\frac{1}{\rho} \left(\Lambda - \frac{qr^*}{2} + \alpha \right) h(A,B) - \frac{(\mu - \alpha)}{\rho} \eta(A) \eta(B). \tag{3.4}$$

By using (2.13) and (2.14) in (3.4), we obtain

$$Ric(A,B) = \left[-\frac{1}{\rho} \left(\Lambda - \frac{qr}{2} - \frac{q\alpha^2(n-1)^2}{2} + \alpha \right) - \alpha^2(n-2) \right] h(A,B) - \left(\frac{(\mu - \alpha)}{\rho} + \alpha^2 \right) \eta(A)\eta(B), \tag{3.5}$$

that is,

$$Ric(A, B) = \sigma_1 h(A, B) + \sigma_2 \eta(A) \eta(B), \tag{3.6}$$

where

$$\sigma_1 = -\frac{1}{\rho} \left(\Lambda - \frac{qr}{2} - \frac{q\alpha^2(n-1)^2}{2} + \alpha \right) - \alpha^2(n-2), \quad \sigma_2 = -\left(\frac{(\mu - \alpha)}{\rho} + \alpha^2\right).$$

Now, if we fix $B = \zeta$ in (3.6), then we can easily get the following relation:

$$Ric(A,\zeta) = \left[-\frac{1}{\rho} \left(\Lambda - \frac{qr}{2} - \frac{q\alpha^2(n-1)^2}{2} + \mu \right) - \alpha^2(n-1) \right] \eta(A). \tag{3.7}$$

Using (2.12) and values of σ_1 and σ_2 in (3.7), we can have (3.2). Also, on contracting (3.6) and using the values of σ_1 and σ_2 , we find

$$r = (n-1)\left(\frac{\mu}{\rho} - \frac{\alpha}{\rho} - \alpha^2(n-1)\right),\tag{3.8}$$



where μ and $\rho(\neq 0)$ are constant.

Thus, (3.6) together with (3.2) and (3.8) give the relation of Λ and μ , which shows that *- η -RYS on α -CM is an η -EM.

Remark 3.2. For the particular value of $\rho = 0$ in (3.1), an n-dimensional α -CM M endowed with $*-\eta$ -RYS $(h, \zeta, \Lambda, \mu, \rho, q)$ furnishes the scalar quantities as $\Lambda = -\alpha + \frac{qr^*}{2}$ and $\mu = \alpha$.

First we give the more general construction of α -cosymplectic manifold:

Example 3.3. Let (N, J, \tilde{h}) be a Kähler manifold. Denote by $\mathbb{R} \times_{\sigma} N$ the manifold $(\mathbb{R} \times N, \Phi, \zeta, \eta, h)$, where Φ is the tensor field such that

$$\begin{split} &\Phi\left(\frac{d}{dt}\right)=0, \quad \Phi(A)=J(A), \quad A\in TN,\\ &\zeta=\frac{d}{dt}, \quad \eta=dt, \quad h=dt\otimes dt+\exp(2\alpha t)\tilde{h}, \quad \alpha\in\mathbb{R}. \end{split}$$

Putting $\sigma = \exp(\alpha t)$, h is the warped product metric of the Euclidean metric and \tilde{h} by means of the function σ . Then $\mathbb{R} \times_{\sigma} N$ is α -cosymplectic and (N, \tilde{h}) is a totally umbilical submanifold with mean curvature vector $-\alpha \zeta$. Assume that $\alpha \neq 0$. Applying well-known curvature formulas, one relates the Ricci tensors of N and $\mathbb{R} \times_{\sigma} N$. But here we consider the flat Kähler manifold \mathbb{R}^4 endowed with the canonical complex structure and then the α -cosymplectic manifold $\mathbb{R} \times_{\sigma} \mathbb{R}^4$. If $\alpha = 0$, one has $\sigma = 1$, $\mathbb{R} \times_{\sigma} N = \mathbb{R} \times N$ is cosymplectic and N is totally geodetic. In this case the Ricci tensors are related by:

$$Ric(A, B) = \tilde{Ric}(A - \eta(A)\zeta, B - \eta(B)\zeta). \tag{3.9}$$

It follows that if N is an Einstein manifold, then $\mathbb{R} \times N$ is η -Einstein.

Next, by giving the following example we can show the existence of this soliton in α -cosymplectic manifold:

Example 3.4. Recall an example of 5-dimensional α -CM in [11], that is,

$$M = \{(x_1, x_2, y_1, y_2, z) \in \mathbb{R}^5, \Phi, \zeta, \eta, h\},\$$

where (x_1, x_2, y_1, y_2, z) are the standard coordinates in \mathbb{R}^5 .

The linearly independent vector fields on M are denoted by $E_1 = \exp^{\alpha z} \partial x_1$, $E_2 = \exp^{\alpha z} \partial x_2$, $E_3 = \exp^{\alpha z} \partial y_1$, $E_4 = \exp^{\alpha z} \partial y_2$ and $E_5 = \zeta = -\partial z$ for $i = \{1, 2\}$. Thus, h and Φ are respectively defined as

$$h(E_i, E_i) = 1$$
, $h(E_i, E_j) = 0$, $i \neq j = \{1, 2, 3, 4, 5\}$



and

$$\Phi E_1 = -E_2, \quad \Phi E_2 = E_1, \quad \Phi E_3 = -E_4, \quad \Phi E_4 = E_3, \quad \Phi E_5 = \Phi \zeta = 0.$$

By the linearity of these tensors, it is quite easy to compute (2.1)-(2.4). Also, (2.6) and (2.7) are verified in [11].

By applying Koszul's formula, Rim of M (see [11]) can be obtained easily and hence the components Ric of Ricci tensor of M are: $Ric(E_i, E_i) = -4\alpha^2$ for $i = \{1, 2, 3, 4, 5\}$. Since $r = \sum_{i=1}^{5} Ric(E_i, E_i)$, so we have $r = -20\alpha^2$.

Now, we use (3.7) and find

$$Ric(E_5, E_5) = Ric(\zeta, \zeta) = \left[-\frac{1}{\rho} \left(\Lambda + 2q\alpha^2 + \mu \right) - 4\alpha^2 \right].$$

By equating the values of $Ric(\zeta,\zeta)$, we arrive at a relation: $\Lambda + \mu = -2q\alpha^2$. We also verify this relation for n=5 by using (3.2). Thus, h gives an *- η -RYS $(h,\zeta,\Lambda,\mu,\rho,q)$ on an α -cosymplectic manifold M of dimension 5.

On the other hand, suppose that an *n*-dimensional α -CM M admitting *- η -RYS $(h, \zeta, \Lambda, \mu, \rho, q)$ satisfies

$$Q(h, Ric)(A, B, C, D) = 0, (3.10)$$

where Q(h, Ric)(A, B, C, D) = (h(A, B).Ric)(C, D), for all vector fields A, B, C, D on M. This can be expressed as

$$Q(h, Ric)(A, B, C, D) = h(B, C)Ric(A, D) - h(A, C)Ric(B, D)$$

$$+ h(B, D)Ric(A, C) - h(A, D)Ric(B, C).$$
(3.11)

Theorem 3.5. If *- η -RYS on an α -CM M satisfies Q(h, Ric) = 0, then

$$\Lambda = \frac{q}{2} \left(r + \alpha^2 (n - 1)^2 \right) - \alpha \left(1 - \alpha \rho \right), \tag{3.12}$$

$$\mu = \alpha (1 - \alpha \rho). \tag{3.13}$$

Proof. From the expressions (3.6), (3.10) and (3.11), we have

$$\sigma_2[h(B,C)\eta(A)\eta(D) - h(A,C)\eta(B)\eta(D) + h(B,D)\eta(A)\eta(C) - h(A,D)\eta(B)\eta(C)] = 0. \quad (3.14)$$

Above equation follows that $\sigma_2 = 0$, which implies that

$$\mu = \alpha(1 - \alpha \rho).$$



We obtain the following from (3.2)

$$\Lambda = \frac{q}{2} \left(r + \alpha^2 (n - 1)^2 \right) - \alpha \left(1 - \alpha \rho \right). \tag{3.15}$$

Now, by using these values of σ_1 , σ_2 and Λ as well as μ in (3.6), we calculate

$$Ric(A, B) = -(\alpha^2(n-1))h(A, B).$$
 (3.16)

Thus, from above we can state the following result:

Corollary 3.6. If *- η -RYS on an α -CM M satisfies Q(h, Ric) = 0, then M is an EM.

Next, we have

$$Rim(\zeta, A).Ric = 0, (3.17)$$

then we have

$$Ric(Rim(\zeta, A)B, C) + Ric(B, Rim(\zeta, A)C) = 0, \tag{3.18}$$

for all vector fields A, B, C on M.

Theorem 3.7. If *- η -RYS on an α -CM M satisfies $Rim(\zeta, A).Ric = 0$, then either M becomes CM or we have

$$\Lambda = \frac{q}{2} \left(r + \alpha^2 (n - 1)^2 \right) - \alpha (1 - \alpha \rho) \tag{3.19}$$

$$\mu = \alpha (1 - \alpha \rho). \tag{3.20}$$

Proof. In view of (3.6) and (3.18), we compute

$$\alpha^2 \sigma_2(2\eta(A)\eta(B)\eta(C) - \eta(C)h(A, B) - \eta(B)h(A, C)) = 0.$$
(3.21)

Putting $C = \zeta$ into (3.21) and using (2.4), it is quite easy to see

$$\alpha^2 \sigma_2 h(\Phi A, \Phi B) = 0, \tag{3.22}$$

which implies either $\alpha = 0$ or $\sigma_2 = 0$. Further, from later case we find $\mu = \alpha(1 - \alpha\rho)$ and hence from (3.2), we calculate the value of Λ . From the first case we can also say that M is CM. This is the desired result.

Next, by using the values of Λ as well μ in (3.6), we have

$$Ric(A, B) = -(\alpha^2(n-1))h(A, B).$$
 (3.23)



Thus, we can state the following:

Corollary 3.8. If $*-\eta$ -RYS on an α -CM M satisfies $Rim(\zeta, A).Ric = 0$ then M is either an EM or CM.

The non-flat manifold M of n-dimension is named pseudo Ricci symmetric, if $Ric \neq 0$ of M satisfies the condition:

$$(\nabla_C Ric)(A, B) = 2\kappa(C)Ric(A, B) + \kappa(A)Ric(C, B) + \kappa(B)Ric(C, A), \tag{3.24}$$

where κ is a non-zero 1-form. In particular, M is said to be Ricci symmetric if $\kappa = 0$.

Theorem 3.9. If an α -CM M admitting *- η -RYS is pseudo-Ricci-symmetric, then M is either Ricci symmetric or CM.

Proof. The covariant derivative of (3.6) leads

$$(\nabla_C Ric)(A, B) = \nabla_C \left[\sigma_1 h(A, B) + \sigma_2 \eta(A) \eta(B)\right] = \alpha \sigma_2 \left[h(\Phi A, \Phi C) \eta(B) + \eta(A) h(\Phi B, \Phi C)\right]. \tag{3.25}$$

Further, we use the relations (3.6), (3.24), (3.25) and obtain

$$2\kappa(C)\left[\sigma_1 h(A,B) + \sigma_2 \eta(A)\eta(B)\right] + \kappa(A)\left[\sigma_1 h(C,B) + \sigma_2 \eta(C)\eta(B)\right]$$

$$+ \kappa(B)\left[\sigma_1 h(C,A) + \sigma_2 \eta(C)\eta(A)\right] = \alpha\sigma_2\left[h(\Phi A, \Phi C)\eta(B) + \eta(A)h(\Phi B, \Phi C)\right].$$
(3.26)

Taking $C = B = \zeta$ in (3.26), we get

$$(\sigma_1 + \sigma_2)(\kappa(A) + 3\eta(A)\kappa(\zeta)) = 0,$$

which gives either

$$\kappa(A) = -3\eta(A)\kappa(\zeta) \tag{3.27}$$

or

$$\sigma_1 + \sigma_2 = 0. \tag{3.28}$$

Putting $A = \zeta$ in (3.27), we have $\kappa(\zeta) = 0$, which further implies that $\kappa(A) = 0$. Also, from (3.28) and (3.2), we can have $\alpha^2(n-1) = 0$. This implies that $\alpha = 0$ because $n \neq 1$. Thus, we arrive at our desired result.



4 Some curvature properties on α -CM admitting *- η -RYS

This section deals with the curvature properties on M admitting *- η -RYS. We mainly discuss the conditions that M is ζ -concircularly flat, ζ -conharmonically flat, ζ - W_2 flat and ζ -conformal flat.

Theorem 4.1. Let M be an n-dimensional α -CM admitting *- η -RYS $(h, \zeta, \Lambda, \mu, \rho, q)$, where ζ being the Reeb vector field on M. Then M is

- (1) ζ -concircularly flat if and only if $\mu = \alpha \rho \alpha^2$.
- (2) ζ -conformal curvature flat.
- (3) ζ -conharmonically flat if and only if $\mu = \alpha + (n-1)\alpha^2 \rho$.
- (4) ζW_2 -curvature flat if and only if $\mu = \alpha \rho \alpha^2$.

Proof. By using the property h(QA, B) = Ric(A, B) in (3.6), we arrive at

$$QB = \sigma_1 B + \sigma_2 \eta(B) \zeta, \tag{4.1}$$

where
$$\sigma_1 = -\frac{1}{\rho} \left(\Lambda - \frac{qr}{2} - \frac{q\alpha^2(n-1)^2}{2} + \alpha \right) - \alpha^2(n-2)$$
 and $\sigma_2 = -\left(\frac{\mu - \alpha}{\rho} + \alpha^2 \right)$.

Firstly, we put $C = \zeta$ into (2.16) and use the relations (2.4), (2.8) and (3.8), we have

$$\overline{\mathbb{C}}(A,B)\zeta = \frac{1}{n} \left(\frac{\mu}{\rho} - \frac{\alpha}{\rho} + \alpha^2 \right) (\eta(A)B - \eta(B)A), \qquad (4.2)$$

which gives $\overline{\mathbb{C}}(A, B)\zeta = 0$ if and only if $\mu = \alpha - \rho\alpha^2$.

Secondly, if we put $C = \zeta$ and use (2.8), (2.12), (4.1) in (2.19), then we have

$$\mathbb{C}^*(A,B)\zeta = \left(\frac{\alpha^2 - \sigma_1}{n-2} + \frac{r}{(n-1)(n-2)}\right)(\eta(B)A - \eta(A)B). \tag{4.3}$$

Again, using the value of σ_1 , (3.2) and (3.8), we have

$$\mathbb{C}^*(A,B)\zeta = 0. \tag{4.4}$$

Thirdly, we take $C = \zeta$ in (2.17) and make use of (2.8), (4.1) and (2.12), we get

$$H(A,B)\zeta = \left(\frac{\sigma_1 - \alpha^2}{n-2}\right)(\eta(A)B - \eta(B)A). \tag{4.5}$$

This implies

$$H(A,B)\zeta = 0$$



if and only if

$$\sigma_1 = \alpha^2$$
.

Thus,

$$H(A,B)\zeta = 0$$

if and only if

$$\mu = \alpha + (n-1)\alpha^2 \rho.$$

Lastly, by taking $C = \zeta$ and using (2.8) and (4.1) in (2.18), we conclude

$$W_2(A,B)\zeta = \left(\alpha^2 + \frac{\sigma_1}{n-1}\right) \left(\eta(A)B - \eta(B)A\right). \tag{4.6}$$

From (4.6),

$$W_2(A, B)\zeta = 0$$

if and only if

$$\alpha^2 + \frac{\sigma_1}{n-1} = 0.$$

This further implies that

$$W_2(A,B)\zeta = 0$$

if and only if

$$\mu = \alpha - \rho \alpha^2$$
.

Remark 4.2. We observe that above results are true only for α -Kenmotsu manifolds because Λ and μ are depending on α . But for $\alpha = 0$, one puts in (3.1) $B = \xi$ obtains

$$\Lambda + \mu = \frac{1}{2}qr.$$

Then (3.1) implies

$$Ric = \frac{\mu}{\rho}(h - \eta \otimes \eta).$$

So, according to the cases μ is zero or non-zero, M is Ricci-flat or η -Einstein for $\alpha = 0$. This is consistent with the formula (3.9), when N is Einstein.

Remark 4.3. If M is a cosymplectic manifold, then we have

$$\overline{\mathbb{C}}(A, B)\zeta = -\left(\frac{\mu}{n-\rho}\right)(\eta(B)A - \eta(A)B).$$

and similar relations can be obtained for $H(A, B)\zeta$ and $W_2(A, B)\zeta$, while

$$\mathbb{C}^*(A,B)\zeta = 0.$$



By the above formulas, one has $\mu = 0$ if and only if $\overline{\mathbb{C}}(A, B)\zeta = 0$ if and only if $H(A, B)\zeta = 0$ if and only if $W_2(A, B)\zeta = 0$.

Acknowledgments

The authors would like to thank the reviewer for the valuable comments and constructive suggestions.

Conflict of Interest: The authors declare no competing interests.

Funding: Not Applicable.

Data Availability: Not Applicable.

Ethical Conduct: The manuscript is not currently being submitted for publication elsewhere

and has not been previously publish.



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